

# Lecture 13 Heat Capacities of Solids

23 Oct 2008

$C_v$  is calculated  $C_v = \frac{\partial E}{\partial T}$

$C_p$  is measured and subject to atmospheric pressure

$C_p - C_v =$  (exact expression at end of Ch. 5 Mandl)

## $C_p$ Measurements

At room T, most solids follow  $C_p \approx 3nNk$

where  $n = \# \text{ atoms per molecule}$

$N = \# \text{ molecules in the solid}$

$k = \text{Boltzmann Constant}$

Dulong - Petit's Rule: For a monatomic molecule,  $n=1$ ,  
for 1 mole  $C_p \approx 3Nk = 3R = 24.93 \text{ J K}^{-1} \text{ mol}^{-1}$

## p. 148 Mandl Table

Follows Rule

Element	$C_p$
Al	24.4
Au	25.4
Cu	24.5

Exceptions

Element	$C_p$
S	22.7
Si	19.9
C(diamond)	6.1

The classical limit of vibrational energy explains this rule.

In non-metals, there are 3 independent vibrations:

$$E_{\text{vib}} = 3NkT = \text{vibration energy of } N \text{ atoms}$$

$$C_v^{\text{vib}} = \frac{\partial E_{\text{vib}}}{\partial T} = 3Nk = 3nNk$$

Diamond is an exception to the rule.

Temperature dependence is not explained by this rule.

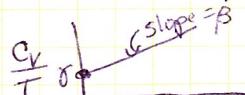
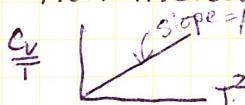
$$\lim_{T \rightarrow 0} C_v = 0$$

$$C_v \propto T^3 \quad C_p \propto T^3 \quad \text{for non-metals}$$

$$C_p = \beta T^3 \quad \text{non-metals}$$

$$C_p = \beta T^3 + \gamma T \quad \text{metals}$$

p. 149



We are not considering heat capacity of conduction electrons because we need to have the fermi-dirac statistics from the end of the course.

# Calculation of Vibrational Heat Capacity

- The number of degrees of freedom are the number of vibrational states
- Theorem of Classical Mechanics

For any system with proper 'normal' coordinates, you can write the vibrational energy as the sum of simple, small angle, harmonic oscillator energies

$$E_{\text{vib}} = \sum_{\alpha=1}^{3N} \left( \frac{p_\alpha^2}{2m_\alpha} + \frac{m_\alpha \omega_\alpha^2 q_\alpha^2}{2} \right)$$

$p_\alpha$  - canonical 'normal' momentum

$q_\alpha$  - normal coordinate

$$p_\alpha = \frac{\partial \mathcal{L}}{\partial \dot{q}_\alpha} \quad \mathcal{L} = \text{Lagrangian}$$

Classically, the amplitude can be anything, the  $E_{\text{cas}}$  become infinite.

- Energy of radiation in cavity is also a sum but over infinite degrees of freedom of the electromagnetic field.
- Quantum mechanics energy of one simple harmonic oscillator

$$E_\alpha = \hbar \omega_\alpha (n_\alpha + \frac{1}{2}) \quad \hbar \omega \text{ is very small}$$

$$E = \lim_{N \rightarrow \infty} \sum_{\alpha=1}^{3N} E_\alpha = \lim_{N \rightarrow \infty} \sum_{\alpha=1}^{3N} \hbar \omega_\alpha (n_\alpha + \frac{1}{2})$$

To Solve:

- (A) Apply partition function to system and heat bath.

System - one harmonic oscillator

heat bath - the rest of the oscillators

$\omega_\alpha$  one mode  
3N normal modes

- (B) Canonical Partition Function:  $Z = \sum_r e^{-E_r/kT}$

- (C) Probability of a particular microstate:  $p_r = \frac{e^{-E_r/kT}}{Z}$

- (D) Non-degenerate quantum state:

$$Z_\alpha = \sum_{n_\alpha=0}^{\infty} e^{-\hbar \omega_\alpha (n_\alpha + \frac{1}{2})/kT} = e^{-\hbar \omega_\alpha / 2kT} [1 + e^{-\hbar \omega_\alpha / kT} + e^{-2\hbar \omega_\alpha / kT} + \dots]$$

$$e^{-\hbar \omega_\alpha / kT} (1 + x + x^2 + \dots) = \left(\frac{1}{1-x}\right) e^{-\hbar \omega_\alpha / 2kT}$$

trick to calculate the sum!

$$S = 1 + x + x^2 + \dots$$

$$Sx = x + x^2 + x^3 + \dots$$

$$S - Sx = 1 \rightarrow S = \frac{1}{1-x} \quad \text{only works when } x < 1$$

## Lecture 13 continued... Heat Capacities of Solids

Partition function is:

$$Z_\alpha = \frac{e^{-\beta \hbar \omega_\alpha / 2}}{1 - e^{-\beta \hbar \omega_\alpha}}$$

- (E) Can find the average energy of a SHO in equilibrium with a heat bath from the partition function.

$$\bar{E}_\alpha = -\frac{\partial \ln Z_\alpha}{\partial \beta} = -\frac{1}{Z_\alpha} \frac{\partial Z_\alpha}{\partial \beta}$$

$$\frac{\partial Z_\alpha}{\partial \beta} = e^{-\beta \hbar \omega_\alpha / 2} \frac{(-\hbar \omega_\alpha e^{-\beta \hbar \omega_\alpha})}{(1 - e^{-\beta \hbar \omega_\alpha})^2} + \frac{e^{-\beta \hbar \omega_\alpha / 2} (-\frac{1}{2} \hbar \omega_\alpha)}{(1 - e^{-\beta \hbar \omega_\alpha})}$$

$$\frac{1}{Z_\alpha} = (1 - e^{-\beta \hbar \omega_\alpha}) e^{+\beta \hbar \omega_\alpha / 2}$$

$$\bar{E}_\alpha = -\frac{1}{Z_\alpha} \frac{\partial Z_\alpha}{\partial \beta} = \frac{\hbar \omega_\alpha e^{-\beta \hbar \omega_\alpha}}{(1 - e^{-\beta \hbar \omega_\alpha})} + \frac{1}{2} \hbar \omega_\alpha$$

$$\boxed{\bar{E}_\alpha = \frac{\hbar \omega_\alpha}{(e^{\beta \hbar \omega_\alpha} - 1)} + \frac{1}{2} \hbar \omega_\alpha}$$

\* This is an extremely important relation to know

↑  
denominator  
depends on T

↑  
no temperature dependence  
never non-zero energy  
but this term doesn't contribute to heat capacity

- (F) Can obtain the DuLroy-Petit Rule from the classical limit

(can Neglect the zero-point energy)

$$\text{But } \lim_{k \rightarrow 0} \bar{E}_\alpha = \lim_{k \rightarrow 0} \left\{ \frac{\hbar \omega_\alpha}{e^{\beta \hbar \omega_\alpha} - 1} + \frac{\hbar \omega_\alpha}{2} \right\} = \lim_{k \rightarrow 0} \left( \frac{kT \left( \frac{\hbar \omega_\alpha}{kT} \right)}{e^{(\hbar \omega_\alpha / kT)} - 1} \right)$$

$$\lim_{k \rightarrow 0} \bar{E}_\alpha = kT \left( \frac{x}{e^x - 1} \right) \quad \text{where } x \ll 1 \text{ at any temperature} \\ e^x \sim 1 + x$$

$$\lim_{k \rightarrow 0} \bar{E}_\alpha \approx kT \left( \frac{x}{1+x-1} \right) = kT$$

If it is classical, there are 3 oscillators  $\bar{E}_\alpha = 3kT$  Dulroy-Petit Rule  
Deviations imply that  $\frac{\hbar \omega}{kT} \ll 1$

(G) Exact expression

$$\bar{E} = \sum_{\alpha=1}^{3N} \bar{E}_{\alpha} = \sum_{\alpha=1}^{3N} \left( \frac{\hbar \omega_{\alpha}}{e^{\hbar \omega_{\alpha}/kT} - 1} + \frac{1}{2} \hbar \omega_{\alpha} \right)$$

(H) Can get the heat capacity from  $\bar{E}$

$$C_V = \left( \frac{\partial \bar{E}}{\partial T} \right)_V = \sum_{\alpha=1}^{3N} k \left( \frac{\hbar \omega_{\alpha}}{kT} \right)^2 \frac{e^{\hbar \omega_{\alpha}/kT}}{(e^{\hbar \omega_{\alpha}/kT} - 1)^2}$$

$$C_V = k \sum_{\alpha=1}^{3N} \left( \frac{\hbar \omega_{\alpha}}{kT} \right)^2 \frac{e^{\hbar \omega_{\alpha}/kT}}{(e^{\hbar \omega_{\alpha}/kT} - 1)^2}$$

↑      ↑      dimensionless terms in the sum

heat capacity is energy per unit temperature

(I) Classical result, simplification of the  $C_V$  sum

If  $\frac{\hbar \omega_{\alpha}}{kT} \ll 1$  for all  $\alpha$ 's

$$C_V = k \sum_{\alpha=1}^{3N} \frac{x_{\alpha}^2 e^{x_{\alpha}}}{(e^{x_{\alpha}} - 1)^2} = k \sum_{\alpha=1}^{3N} 1 = 3Nk \quad \begin{matrix} \text{again Dulong-Petit rule} \\ \text{by SHO when } \hbar \omega_{\alpha}/kT \ll 1 \end{matrix}$$

Aluminum - works @ Room Temperature    Carbon - deviates because  $\hbar \omega_{\alpha}/kT$  is not  $\ll 1$

(J) Einstein's Model

It is a formidable task to find the  $\omega_{\alpha}$ 's for  $3N$  unknowns.

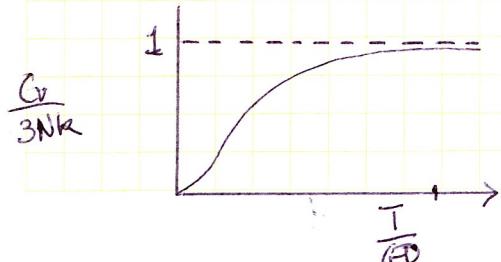
Einstein says  $\omega_{\alpha} = \omega_E$  for all  $\alpha$

$$C_V = k (3N) \frac{(\hbar \omega_E)^2}{(kT)} e^{\hbar \omega_E/kT} \frac{1}{(e^{\hbar \omega_E/kT} - 1)^2}$$

Einstein was able to calculate  $\omega_E$  for an elastic solid.

(K)  $\omega_E = \frac{\hbar \omega_E}{k}$  = Einstein temperature, a characteristic of the solid.

$$C_V = \frac{3Nk \left( \frac{\omega_E}{T} \right)^2 e^{\omega_E/T}}{(e^{\omega_E/T} - 1)^2} \quad \begin{matrix} \lim C_V = 0 & \lim C_V = 3Nk \text{ approaches} \\ T \rightarrow 0 & T \rightarrow \infty \text{ asymptotically} \end{matrix}$$



For one mole of solid  $\frac{C_V}{3R}$

Deviations  $C_V \ll 3R$  for Carbon